### DISSIPATION AND AMPLIFICATION OF GENERALIZED GAUSSIAN STATES1

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solution for the density matrix elements is given by the same function as initially, with the only difference that the field parameters become time-dependent. two generalizations of this GSF. The final result shows that the time-dependent to the Gaussian state field (GSF) (a field with Gaussian Wigner function) and to of damping and amplification. We apply these results to three different fields, i.e., define a generating function of this field and find its time-development in the case We study the damping and amplification of a single-mode radiation field. We

## 1. Master equations and their solutions

where  $H_0 = \hbar\omega(a^+a + \frac{1}{2})$ ,  $\omega$  is the resonant radiation frequency, and a and  $a^+$  are the equation [1] reservoir in the Born-Markov approximation can be described by the following master The dynamics of single-mode radiation field, interacting with a zero-temperature  $\frac{\partial \rho}{\partial t} = \frac{1}{i\hbar} [H_0, \rho] + \frac{\gamma}{2} (2a\rho a^+ - a^+ a\rho - \rho a^+ a),$ 

$$\frac{\partial \rho}{\partial t} = \frac{1}{i\hbar} [H_0, \rho] + \frac{I}{2} (2a\rho a^+ - a^+ a \rho - \rho a^+ a),$$
  
 $\frac{1}{2} + \frac{1}{2} [H_0, \rho] + \frac{I}{2} (2a\rho a^+ - a^+ a \rho - \rho a^+ a),$ 

(1)

master equation for the perfect amplifier is given by annihilation and creation operators and  $\gamma$  is the damping constant. Analogously, the

$$\frac{\partial \rho}{\partial t} = \frac{1}{i\hbar} [H_0, \rho] + \frac{\gamma}{2} (2a^+\rho a - aa^+\rho - \rho aa^+). \tag{2}$$

damping, the differential equation The corresponding density-matrix elements in the Fock basis satisfy, in the case of

$$\frac{\partial}{\partial t} \langle m | \rho(t) | n \rangle = -i\omega(m-n) \langle m | \rho(t) | n \rangle - \frac{\gamma}{2} (m+n) \langle m | \rho(t) | n \rangle + \gamma \sqrt{(n+1)(m+1)} \langle m+1 | \rho(t) | n+1 \rangle.$$
 (3)

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The analogous equation in the case of amplification reads as

$$\frac{\partial}{\partial t} \langle m | \hat{\rho}(t) | n \rangle = -i\omega(m-n) \langle m | \rho(t) | n \rangle - \frac{\gamma}{2} (m+n+2) \langle m | \rho(t) | n \rangle = 10$$

$$+\gamma \sqrt{nm} \langle m-1 | \rho(t) | n-1 \rangle . \tag{4}$$

of Arnoldus [6]. From eqs. (3) and (4), after some calculations, we obtain decoupled The above equations can be solved by using different methods and the solutions are well-known [2-5]. We would like to give a simplified treatment being similar to that amplification, respectively equations for normally and antinormally ordered moments in the case of damping and

$$\frac{d}{dt}m_{k,l}^{N}(t) = [i\omega(k-l) - \frac{\gamma}{2}(k+l)]m_{k,l}^{N}(t),$$
(5)

$$\frac{d}{dt}m_{k,l}^{A}(t) = [i\omega(k-l) + \frac{\gamma}{2}(k+l)]m_{k,l}^{A}(t),$$

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where the normally and antinormally ordered moments are given by

$$m_{k,l}^{N}(t) \equiv \left\langle (a^{+})^{k} a^{l} \right\rangle(t) = Tr[(a^{+})^{k} a^{l} \rho(t)] = \sum_{h=0}^{\infty} \frac{\sqrt{(k+h)!(l+h)!}}{h!} \left\langle l+h | \rho(t) | k+h \right\rangle,$$

$$m_{k,l}^{A}(t) \equiv \left\langle a^{l}(a^{+})^{k} \right\rangle(t) = Tr[a^{l}(a^{+})^{k}\rho(t)] = \sum_{h=0}^{\infty} \frac{(k+l+h)!}{\sqrt{(k+h)!(l+h)!}} \left\langle l+h|\rho(t)|k+h \right\rangle.$$

solution for the damping is In both equations the moments for different k and l evolve independently. The ution for the damping is

$$m_{k,l}^{N}(t) = \Gamma_{N}^{k} \Gamma_{N}^{*l} m_{k,l}^{N}(0),$$
 (9)

$$\Gamma_N(t) \equiv \exp\left[\left(i\omega - \frac{\gamma}{2}\right)t\right]$$
 , so the second of the

and that for the amplification is

$$m_{k,l}^{A}(t) = \Gamma_{A}^{k} \Gamma_{A}^{*l} m_{k,l}^{A}(0)$$
, (11)

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$$\Gamma_A(t) \equiv \exp\left[\left(i\omega + \frac{\gamma}{2}\right)t\right]$$
.

It is convenient to express the density matrix elements by using the generating function. By using the generating function [7]

$$G_s(\lambda_1, \lambda_2, \lambda_0; t) = Tr[\rho(t)\{\exp(\lambda_1 a^+ + \lambda_2 a - \lambda_0 a^+ a)\}_s]$$
(13)

the density matrix elements can be expressed in terms of

which matrix elements can be expressed in terms of
$$\langle m | \rho(t) | n \rangle = Tr(\rho(t) | n \rangle \langle m |)$$

$$= \frac{1}{\sqrt{m! \, n!}} \left( \frac{\partial}{\partial \lambda_1} \right)^n \left( \frac{\partial}{\partial \lambda_2} \right)^m G_N(\lambda_1, \lambda_2, \lambda_0 = 1; t) |_{\lambda_1 = \lambda_2 = 0}$$
(14)

and the ordered moments of annihilation and creation operators in terms of

$$m_{k,l}^{s}(t) = \left\langle \left\{ (a^{+})^{k} a^{l} \right\}_{s} \right\rangle (t) = Tr[\rho(t)\{(a^{+})^{k} a^{l}\}_{s}]$$

$$= \left( \frac{\partial}{\partial \lambda_{1}} \right)^{k} \left( \frac{\partial}{\partial \lambda_{2}} \right)^{l} G_{s}(\lambda_{1}, \lambda_{2}, \lambda_{0} = 0; t)|_{\lambda_{1} = \lambda_{2} = 0} . \tag{15}$$

The subscripts s=1 (or N), s=-1 (or A) and s=0 (or W) in eqs. (13)–(15) denote the operator ordering: normal, antinormal and symmetrical (Weyl).  $\{(a^+)^k a^l\}_W$  denotes the expectation value of the symmetrized form of the operator  $(a^+)^k a^l$  [8]. (k+l)!/(k!l!) ways. The symmetrically (Weyl) ordered product of these operators, denoted by  $\{(a^+)^k a^l\}_W$ , is just the average of all of these differently ordered products. Eq. (14) has been obtained by making use of the relation (see e.g. [1]) This product of k creation operators and l annihilation operators can be ordered in

$$|n\rangle\langle m| = \lim_{\lambda_0 \to 1} \frac{1}{\sqrt{m! \, n!}} \sum_{h=0}^{\infty} \frac{1}{h!} (-\lambda_0)^h (a^+)^{h+n} a^{h+m}$$
 (16)

and employing eq. (13). From eqs. (13)-(15) we obtain directly the Taylor series expansion of the generating function

$$G_s(\lambda_1, \lambda_2, \lambda_0; t) = \sum_{j_1, j_2, j_0 = 0}^{\infty} \frac{\lambda_1^{j_1} \lambda_2^{j_2} (-\lambda_0)^{j_0}}{j_1! j_2! j_0!} m_{j_1 + j_0, j_2 + j_0}^s(t)$$
(17)

or

$$\begin{array}{ll} G_N(\lambda_1,\lambda_2,\lambda_0;t) & = & \sum_{j_1,j_2,j_0=0}^{\infty} \frac{\lambda_1^{j_1}\lambda_2^{j_2}(1-\lambda_0)^{j_0}}{j_1!j_2!j_0!} \\ & \times \sqrt{(j_1+j_0)!(j_2+j_0)!}\langle j_2+j_0|\rho(t)|j_1+j_0 \rangle \ . \end{array}$$

(18)

By employing eqs. (17) and (9) the relation for the damping follows

$$G_N(\lambda_1, \lambda_2, \lambda_0; t) = G_N(\lambda_1', \lambda_2', \lambda_0'; 0), \tag{19}$$

$$\lambda_1'(t) = \Gamma_N(t)\lambda_1, \ \lambda_2'(t) = \Gamma_N^*(t)\lambda_2, \ \lambda_0'(t) = |\Gamma_N(t)|^2 \lambda_0.$$
 (20)

Analogously, by using eq. (11) the relation for the amplification follows

$$G_A(\lambda_1, \lambda_2, \lambda_0; t) = G_A(\lambda_1', \lambda_2', \lambda_0'; 0), \qquad (21)$$

$$\lambda_1'(t) = \Gamma_A(t)\lambda_1, \ \lambda_2'(t) = \Gamma_A^*(t)\lambda_2, \ \lambda_0'(t) = |\Gamma_A(t)|^2 \lambda_0.$$
 (22)

The results for the matrix elements in the case of damping read as

$$\langle m | \rho(t) | n \rangle = \Gamma_N^n(t) \Gamma_N^{\star m}(t) \sum_{h=0}^{\infty} \frac{1}{h!} \sqrt{\frac{(m+h)!(n+h)!}{m!n!}} \times \left(1 - \left|\Gamma_N(t)\right|^2\right)^h \langle m+h| \rho(0) | n+h \rangle , \qquad (23)$$

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where we used eq. (19) along with eq. (14).

The treatment in the amplification was represented by using the normally ordered generating function  $G_N$ . For generating functions of the different order N and A the following relationship holds of The treatment in the amplification case is similar, with the only difference that eq. 1) should be rewritten by using the normally ordered generating function of the eq.

$$G_N(\lambda_1, \lambda_2, \lambda_0; t) = \frac{1}{1 - \lambda_0} \exp\left[-\lambda_1 \lambda_2 \left(1 - \lambda_0\right)^{-1}\right] G_A(\Lambda_1, \Lambda_2, \Lambda_0; t), \tag{24}$$

$$\Lambda_i = rac{\lambda_i}{1-\lambda_0}, \quad i=0,1,2$$
 .

The above equation can be obtained quite analogously as in paper [9], in which the relationship between  $G_s$  and  $G_W$  has been found. By inserting eq. (21) into eq. (24) we obtain

$$G_N(\lambda_1, \lambda_2, \lambda_0; t) = \frac{1}{1 - \lambda_0} \exp\left[-\lambda_1 \lambda_2 (1 - \lambda_0)^{-1}\right] G_A(\Lambda_1', \Lambda_2', \Lambda_0'; 0), \tag{26}$$

$$\Lambda' - \Gamma_{-}(t)\Lambda$$
  $\Lambda' = \Gamma_{+}(t)\Lambda$ 

(27)

 $\Lambda_1'=\Gamma_A(t)\Lambda_1,\quad \Lambda_2'=\Gamma_A^*(t)\Lambda_2,\quad \Lambda_0'=\left|\Gamma_A(t)\right|^2\Lambda_0\,.$  By using eq. (24) we can express  $G_A(0)$  in eq. (26) by

$$G_A(\Lambda_1', \Lambda_2', \Lambda_0'; 0) = \frac{1}{1 - \lambda_0'} \exp\left[\lambda_1' \lambda_2' (1 - \lambda_0')^{-1}\right] G_N(\lambda_1', \lambda_2', \lambda_0'; 0), \tag{28}$$

where  $\lambda_i'$  is defined by

$$\frac{\lambda_i'}{1 - \lambda_0'} = \Lambda_i', \quad i = 0, 1, 2. \tag{29}$$

$$G_N(\lambda_1, \lambda_2 \lambda_0; t) = \frac{1}{E} \exp\left[\lambda_1 \lambda_2 \left( |\Gamma_A(t)|^2 - 1 \right) E^{-1} \right] G_N(\lambda_1', \lambda_2', \lambda_0'; 0), \tag{30}$$

$$\lambda_1'(t) = \frac{\Gamma_A(t)}{E} \lambda_1, \ \lambda_2'(t) = \frac{\Gamma_A^*(t)}{E} \lambda_2, \ \lambda_0'(t) = \frac{|\Gamma_A(t)|^2}{E} \lambda_0, \ E \equiv 1 - \left(1 - |\Gamma_A(t)|^2\right) \lambda_0.$$

In obtaining eq. (31) we made use of eqs. (25), (27) and (29). By applying the Taylor expansion (18) on both sides of eq. (30) (and additionally expanding the exponantial function on the rhs. of eq. (30) as well) and main and ma solutions for the density matrix elements in the case of the amplification function on the rhs. of eq. (30) as well), and using eq. (14) we obtain the known

$$\langle m | \rho(t) | n \rangle = \Gamma_A^{-m-1}(t) \Gamma_A^{*-n-1}(t) \sum_{h=0}^{\min(m,n)} \frac{1}{h!} \sqrt{\frac{m! n!}{(m-h)!(n-h)!}} \times \left( |\Gamma_A(t)|^2 - 1 \right)^h \langle m-h | \rho(0) | n-h \rangle.$$
 (3)

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two different generalisations of the GSF. The initial fields will be defined in the following state field (GSF) (i.e. single-mode radiation field with Gaussian Wigner function) and We apply the results in the preceding section to different initial fields: The Gaussian We can write the initial density operator of the GSF in the form [9]

$$\rho(0) = \left(\tau^2 - 4|\mu_A|^2 - \frac{1}{4}\right)^{-\frac{1}{2}} \exp\left\{-\frac{1}{\sqrt{\tau^2 - 4|\mu_A|^2}} \ln\left(\frac{\sqrt{\tau^2 - 4|\mu_A|^2} + \frac{1}{2}}{\sqrt{\tau^2 - 4|\mu_A|^2} - \frac{1}{2}}\right)\right\}$$

$$\times \left[ \mu_A (a - \alpha_0)^2 + \mu_A^* (a^+ - \alpha_0^*)^2 + \tau (a^+ - \alpha_0^*) (a - \alpha_0) + \frac{\tau}{2} \right] \right\}.$$
 (33)

The independent parameters of the GSF are given by the lower-order moments of annihilation and creation operators a and  $a^+$  at the initial time

$$\langle a \rangle (0) = \alpha_0, \quad \langle a^2 \rangle (0) = -2\mu_A^* + \alpha_0^2, \quad \langle a^+ a \rangle (0) = \tau - \frac{1}{2} + |\alpha_0|^2.$$
 (3)

The initial states of the generalisation of the GSF are defined by the density operators [9]

$$\bar{\rho}(0) = \bar{N} a^L \rho(0) (a^+)^L$$
,

$$\tilde{\rho}(0) = \tilde{\mathcal{N}}(a^+)^M \rho(0) a^M \,,$$

where  $\rho(0)$  is the initial density operator of the GSF and  $\bar{N}$  and  $\tilde{N}$  are the normalization

The generating function of the initial GSF

$$G_s(\lambda_1, \lambda_2, \lambda_0; t = 0) = \Phi(s, \lambda_0) \exp \left[ y_1 \lambda_1 + y_2 \lambda_2 - \frac{1}{2} \sum_{i,j=1}^2 \lambda_i c_{ij} \lambda_j \right],$$
 (37)

can be expressed as a sum over the Hermitian polynomials of two variables [9]

$$G_s(\lambda_1, \lambda_2, \lambda_0; 0) = \Phi(s, \lambda_0) \sum_{m_1 = 0}^{\infty} \sum_{m_2 = 0}^{\infty} \frac{\lambda_1^{m_1}}{m_1!} \frac{\lambda_2^{m_2}}{m_2!} H_{m_1, m_2}^{(c)}(x_1, x_2).$$
 (38)

field parameters  $\alpha_0, \mu_A, \tau$ , the ordering parameter s and  $\lambda_0$ . The function  $\Phi(s,\lambda_0)$  and the variables  $c=(c_{ij}),\ y_1,\ y_2,\ x_1$  and  $x_2$  are functions of the

By applying the derivatives (14) and (15) to the generating function (38) we obtain for the matrix elements and the ordered moments of the GSF the following results [9]:

$$\langle m | \rho(0) | n \rangle = \frac{1}{\sqrt{m! n!}} \Phi_0 H_{n,m}^{(\tilde{c})}(\tilde{x}_1, \tilde{x}_2),$$
 (39)

$$_{0} \equiv \frac{1}{\sqrt{(\tau + \frac{1}{2})^{2} - 4|\mu_{A}|^{2}}} \exp \left[ -\frac{\mu_{A}\alpha_{0}^{2} + \mu_{A}^{*}\alpha_{0}^{*2} + (\tau + \frac{1}{2})|\alpha_{0}|^{2}}{(\tau + \frac{1}{2})^{2} - 4|\mu_{A}|^{2}} \right], \tag{40}$$

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$$\begin{split} m_{k,l}^{s}(0) &= \left\langle \left\{ (a^{+})^{k}a^{l} \right\}_{s} \right\rangle(0) = H_{k,l}^{(c(s))}\left(x_{1}(s), x_{2}(s)\right) \\ &= \left[ \frac{k}{2} \right] \left[ \frac{l}{2} \right]_{h_{m}} \sum_{w=0}^{k} \sum_{k=0}^{k!l!} \frac{k!l!(-1)^{v+w} (\alpha_{0}^{\star})^{k-h-2v} (\mu_{A})^{v} [\tau - \frac{s}{2}]^{h} (\mu_{A}^{\star})^{w} \alpha_{0}^{l-h-2w}}{(k-h-2v)!v!h!w!(l-h-2w)!} \right]_{0.51k}^{4k} \\ h_{m} &= \min(k-2v, l-2w) \,. \end{split}$$

By using the moments of the GSF  $m_{k,l}^s(0)$ , the moments of the generalisation of the GSF (35) and (36) can be written as [9]

$$\bar{m}_{k,l}^{N}(0) = \frac{m_{k+L,l+L}^{N}(0)}{m_{L,L}^{N}(0)}, \tag{42}$$

$$\tilde{m}_{k,l}^{A}(0) = \frac{m_{k+M,l+M}^{A}(0)}{m_{M,M}^{A}(0)}.$$
(43)

ments of the generalisation of the GSF [cf. (35) and (36)] take the form Analogously, by using the density matrix elements of GSF the density matrix ele-

$$\langle m | \bar{\rho}(0) | n \rangle = \sqrt{\frac{(m+L)!(n+L)!}{m!n!}} \frac{\langle m+L | \rho(0) | n+L \rangle}{mL_{,L}(0)},$$
 (44)

$$\langle m | \, \tilde{\rho}(0) \, | n \rangle = \sqrt{\frac{m! n!}{(m-M)! (n-M)!}} \frac{\langle m-M | \, \rho(0) \, | n-M \rangle}{m_{M,M}^{A}(0)}, \quad m, n \ge M,$$

$$\langle m | \, \tilde{\rho}(0) \, | n \rangle = 0, \quad m < M \quad \text{or} \quad n < M.$$
(45)

# 3. Dissipation in the case of different GSF's

moments (9) yield the simple resul-The initial moments of the GSF (given by (41)) inserted into the solution for the

$$m_{k,l}^{N}(t) = m_{k,l}^{N}(0) \Big| \alpha_0 \to \alpha_0(t) ,$$

$$\mu_A \to \mu_A(t)$$

$$\tau \to \tau(t)$$

$$(46)$$

the corresponding time-dependent parameters where the time-independent parameter  $\alpha_0, \mu_A, \tau$  appearing in eq. (41) are replaced by

$$\alpha_0(t) = \Gamma_N^*(t)\alpha_0, \ \mu_A(t) = \Gamma_N^2(t)\mu_A, \ \tau(t) = |\Gamma_N(t)|^2 \left(\tau - \frac{1}{2}\right) + \frac{1}{2}. \tag{47}$$

in the following form Analogously, the time-dependent density matrix-elements of the GSF can be written

$$\langle m | \rho(t) | n \rangle = \langle m | \rho(0) | n \rangle | \alpha_0 \to \alpha_0(t)$$

$$\mu_A \to \mu_A(t)$$

$$\tau \to \tau(t)$$

(48)

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Analogously, for the generalisation of the GSF  $\bar{\rho}$  it follows that

$$\bar{m}_{k,l}^{N}(t) = \frac{m_{k+L,l+L}^{N}(0)}{m_{L,L}^{N}(0)} \begin{vmatrix} \alpha_{0} \to \alpha_{0}(t) \\ \mu_{A} \to \mu_{A}(t) \\ \tau, \to \tau(t) \end{vmatrix} , \tag{49}$$

$$\langle m | \bar{\rho}(t) | n \rangle = \sqrt{\frac{(m+L)!(n+L)!}{m!n!}} \frac{\langle m+L | \rho(0) | n+L \rangle}{m_{L,L}^{N}(0)} \begin{vmatrix} \alpha_0 \to \alpha_0(t) \\ \mu_A \to \mu_A(t) \\ \tau \to \tau(t) \end{vmatrix}, \tag{50}$$

density operator of the GSF. where the time-dependent parameters are again given by eqs. (47) and  $\rho(0)$  is the

# 4. Amplification in the case of different GSF's

moments of the GSF can be obtained: In the case of the amplification by using eqs. (41) and (11) the time-dependent

$$m_{k,l}^{A}(t) = m_{k,l}^{A}(0) \Big|_{\alpha_0 \to \alpha_0(t)} ,$$

$$\mu_A \to \mu_A(t)$$

$$\tau \to \tau(t)$$
(51)

where

$$\alpha_0(t) = \Gamma_A^*(t)\alpha_0, \ \mu_A(t) = \Gamma_A^2(t)\mu_A, \ \tau(t) = |\Gamma_A(t)|^2 \left(\tau + \frac{1}{2}\right) - \frac{1}{2}$$

written in the following form have to be used. Further, the time-dependent matrix-elements of the GSF can be

$$\langle m | \rho(t) | n \rangle = \langle m | \rho(0) | n \rangle | \alpha_0 \to \alpha_0(t) .$$

$$\mu_A \to \mu_A(t)$$

$$\tau \to \tau(t)$$

$$(53)$$

the time-dependent moments and matrix-elements Analogously, in the case of the amplification we get for the generalisation of the GSF  $\tilde{
ho}$ 

$$\tilde{m}_{k,l}^{A}(t) = \frac{m_{k+M,l+M}^{A}(0)}{m_{M,M}^{A}(0)} \begin{vmatrix} \alpha_0 \to \alpha_0(t) \\ \mu_A \to \mu_A(t) \end{vmatrix},$$
(54)

$$\langle m | \tilde{\rho}(t) | n \rangle = \sqrt{\frac{m! n!}{(m-M)! (n-M)!}} \frac{\langle m-M | \rho(0) | n-M \rangle}{m_{M,M}^{A}(0)} \left| \begin{array}{c} \alpha_0 \to \alpha_0(t) \\ \mu_A \to \mu_A(t) \end{array} \right|,$$
 (55)

where the time-dependent parameters are given by eq. (53) and  $\rho(0)$  is the density operator of the GSF

### 5. Conclusions

become time-dependent. by the same function as initially, with the only difference that the field parameters result shows that the time-dependent solution for the density matrix elements is given elements and photon distributions for the GSF and its two generalizations. The final results have been applied for the description of the dynamics of the density matrix tial functions. By using these results the time development of the generating function has been calculated for any initial field state for both above mentioned cases. These moments (normally or antinormally ordered) can be described by appropriate exponenboth cases, dissipation and amplification, the time-development of the corresponding We studied the dissipation and amplification of a single-mode radiation field. In

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### References

- [1] W.H. Louisell: Quantum Statistical Properties of Radiation (John Wiley, New York,
- [2] P. Marian, T.A. Marian: Phys. Rev. A 47 (1993) 4487;
  [3] S.M. Barnett, P.L. Knight: Phys. Rev. A 33 (1986) 2444; S.J.D. Phoenix: Phys. Rev. A 41 (1990) 5132;

M. S. Kim: J. Mod. Opt. 7 (1993) 1331;

- [4] V.Bužek, C.H. Keitel, P.L. Knight: Phys. Rev. A 51 (1995) 2575;
- V.Bužek, M.S. Kim, Min Gyu Kim: J. of Korean Phys. Soc. 28 (1995) 123;

- [5] V. Peřinova, A. Lukš, P. Szlachetka: J. Mod. Opt. 36 (1989) 1435;
  [6] H. F. Arnoldus: J. Mod. Opt. 41 (1994) 503;
  [7] J. Peřina: Quantum Statistics of Linear and Nonlinear Optical Phenomena (D. Reidel Publishing Company, Dordrecht/Boston/Lancaster, 1984);
- $\infty$ G.S. Agarwal, E. Wolf: Phys. Ref. D 2 (1970) 2161; ibid 2187; ibid 2206;
- [9] G. Adam: Phys. Lett. A 171 (1992) 66;
- G. Adam: J. Mod. Opt., in print;